

# ***Mantle Geotherms from Geodynamics***

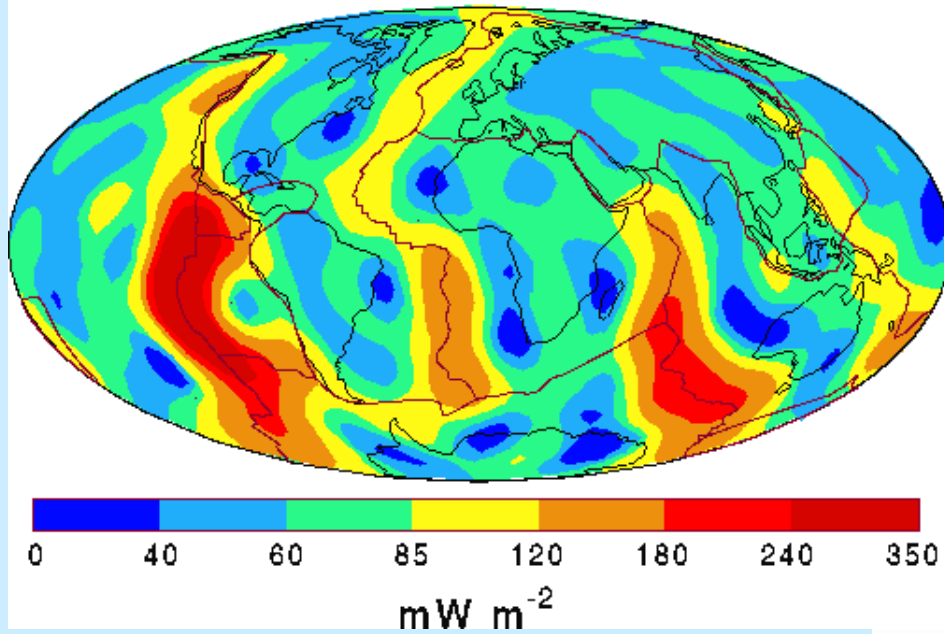
**Shijie Zhong**

**Dept. of Physics, University of Colorado  
Boulder, Colorado, USA**

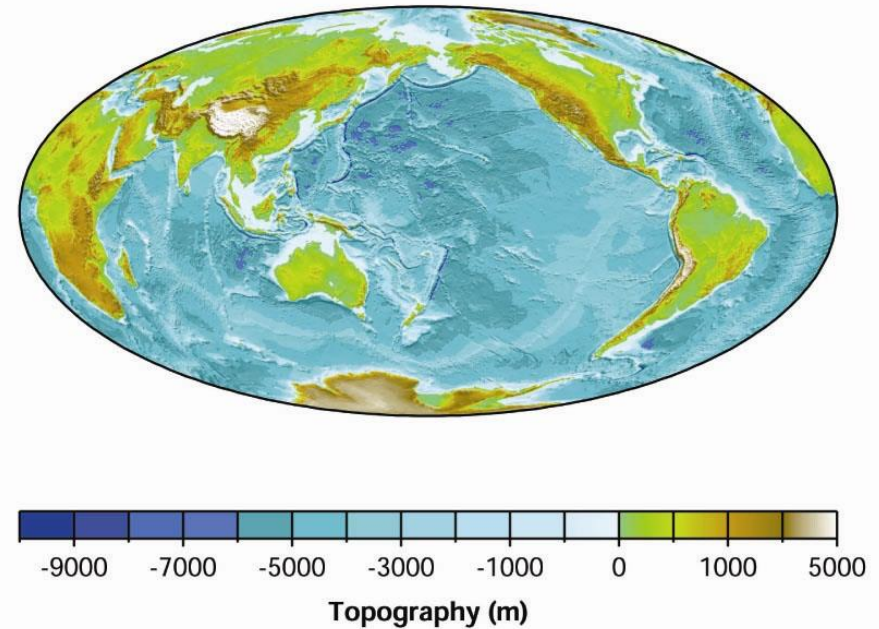
**For CIDER Geotherms workshop**

# Earth's Surface Heat Flux and Topography

Heat Flow



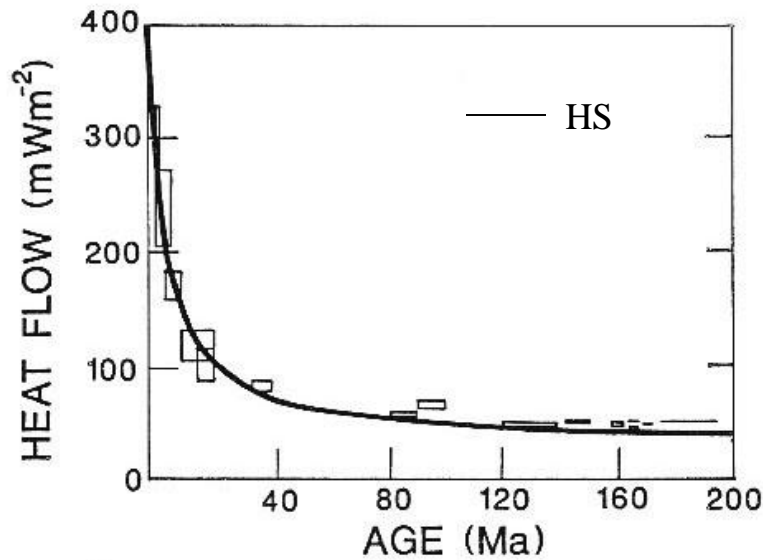
Topography



*Pollack et al., 1993*

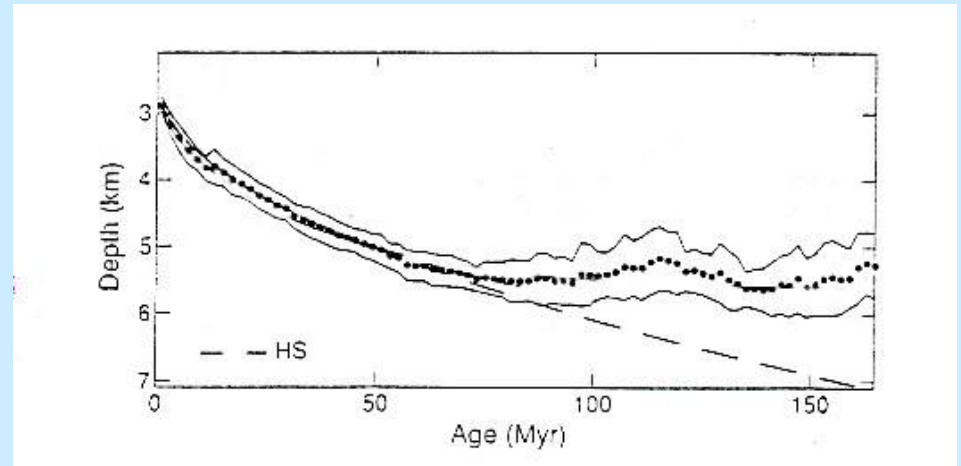
# Age-dependent Heat Flux and Ocean Depth

## Heat Flux



*Lister et al., 1991*

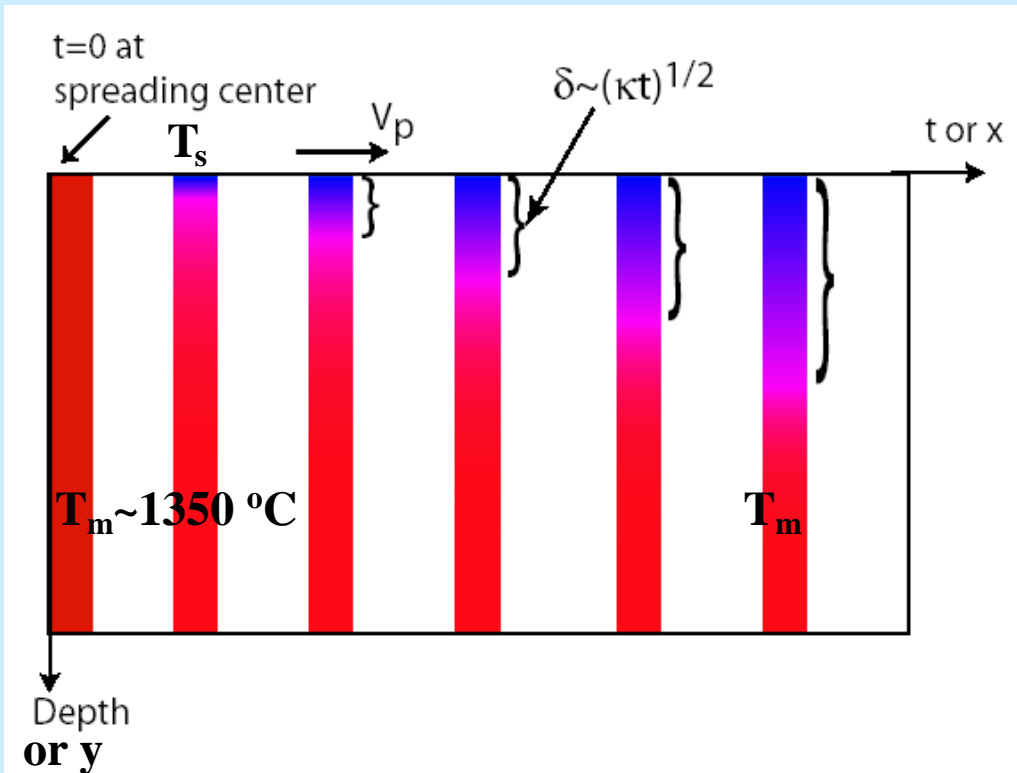
## Ocean Depth



*Stein & Stein, 1992*

# A Simple Model for Age-dependent Topography and Heat Flux (a half-space cooling model)

- *Conductive cooling of oceanic lithosphere as it moves away from the spreading centers [Turcotte & Oxburgh, 1967].*



**Temperature:**

$$T = T_s + (T_m - T_s) \operatorname{erf}[y / (4\kappa t)^{1/2}]$$

**Heat flux:**

$$Q_s \sim (T_m - T_s) / \delta$$

$$\text{or } Q_s = k(T_m - T_s) / (\kappa t)^{1/2}$$

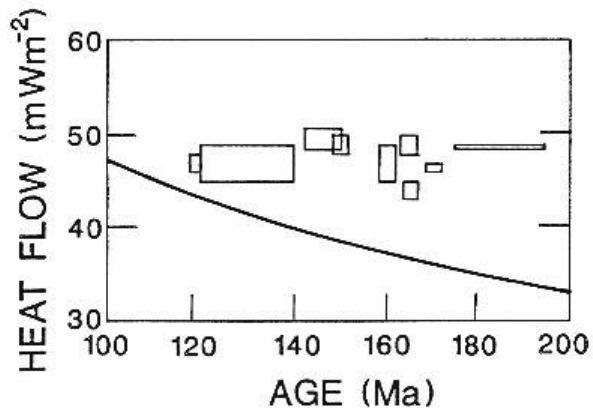
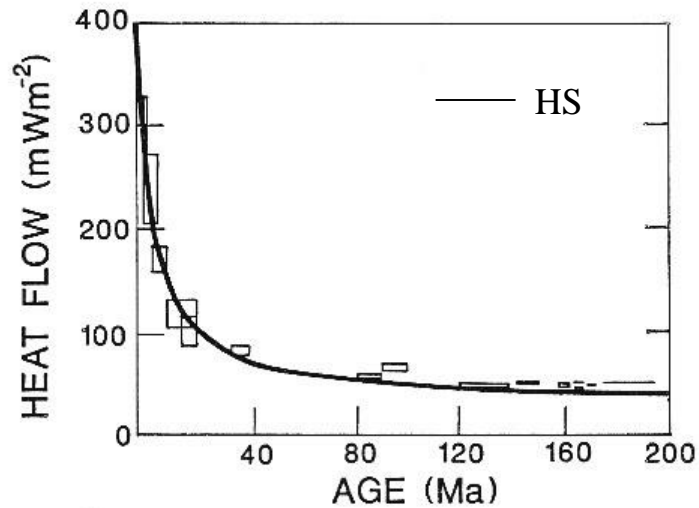
**Topography:**

$$w = 2b\alpha(T_m - T_s)(\kappa t / \pi)^{1/2}$$

$$\text{where } b = \rho_m / (\rho_m - \rho_w)$$

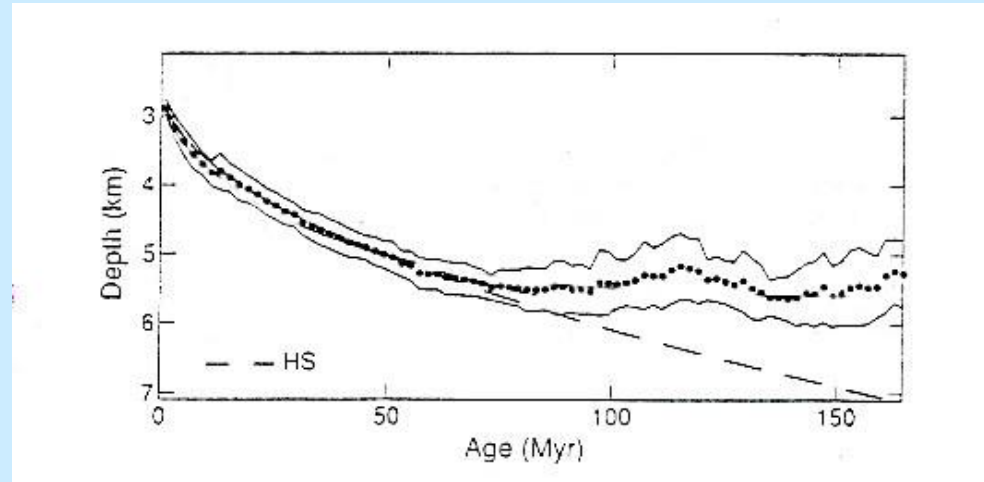
# Deviations from the Half-space Cooling Model at Old Seafloor

## Heat Flux



*Lister et al., 1991*

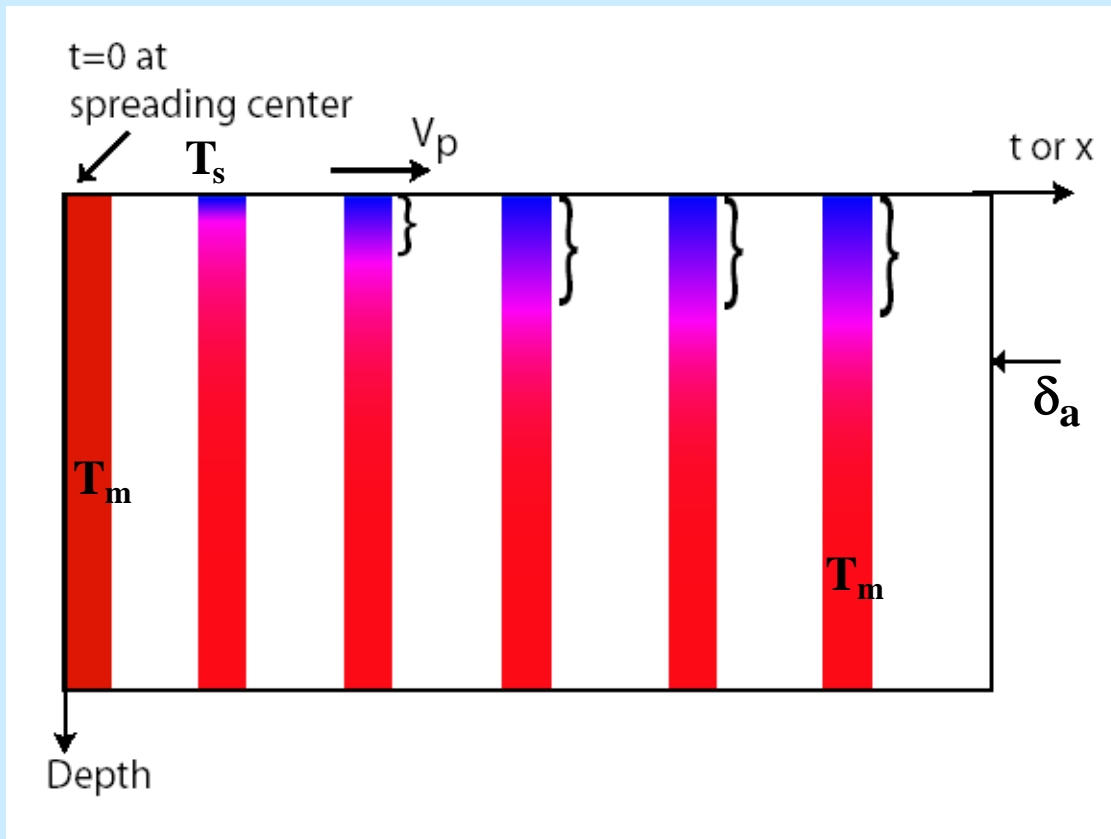
## Ocean Depth



*Stein & Stein, 1992*

# *The Plate Model – a revised model to fit the data*

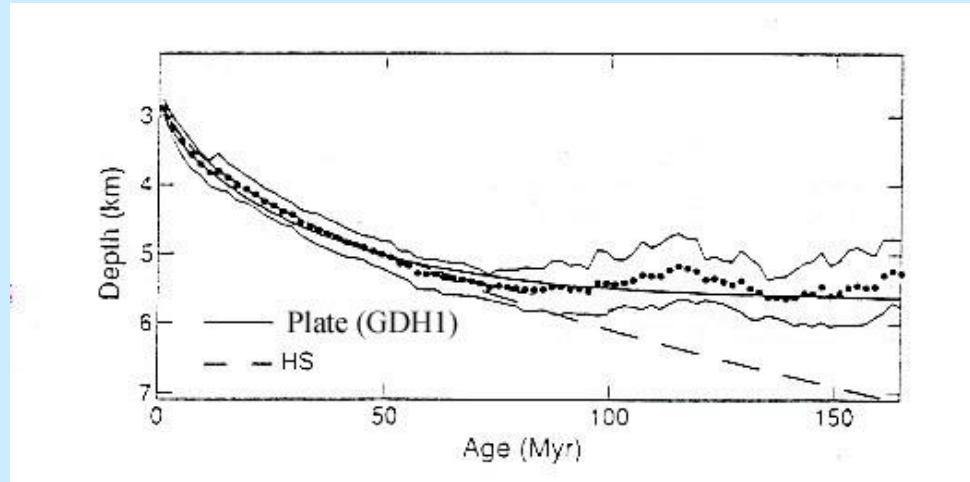
[Parsons & Sclater, 1977; Parsons & McKenzie, 1978]



**There is an upper limit on  $\delta$ ,  $\delta_a$ . And the cooling never reaches to depth of  $\delta_a$ .**

# Improved Fit from the Plate Model

## *Ocean Depth*

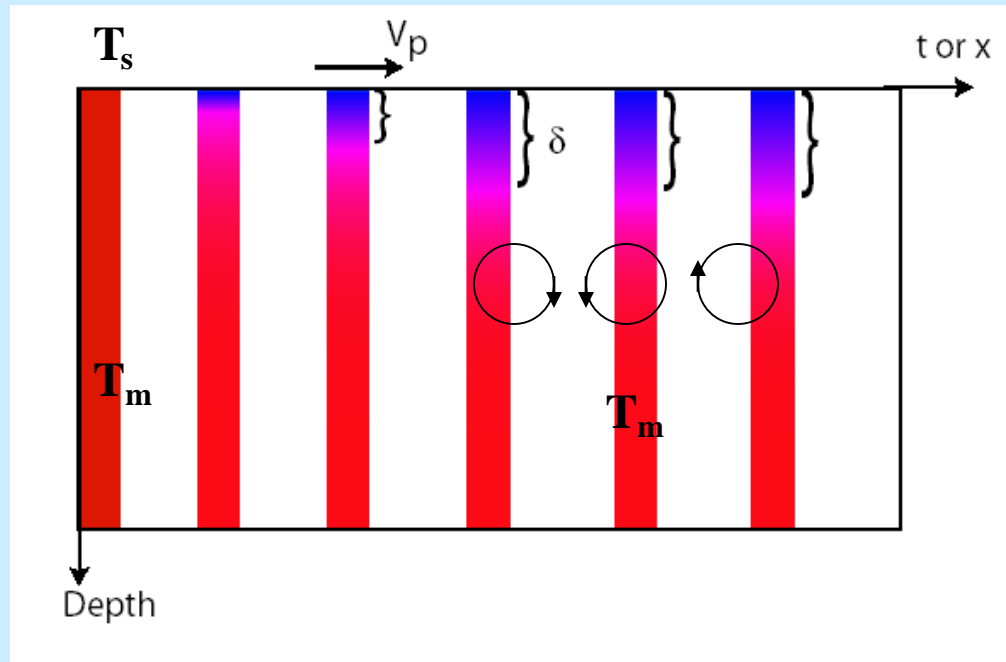


***Stein & Stein, 1992***

***An empirical model!***

# *Small-scale Convection (SSC) Below Old Oceanic Lithosphere -- Physical Basis for the Plate Model*

*[Parsons and McKenzie, 1978]*



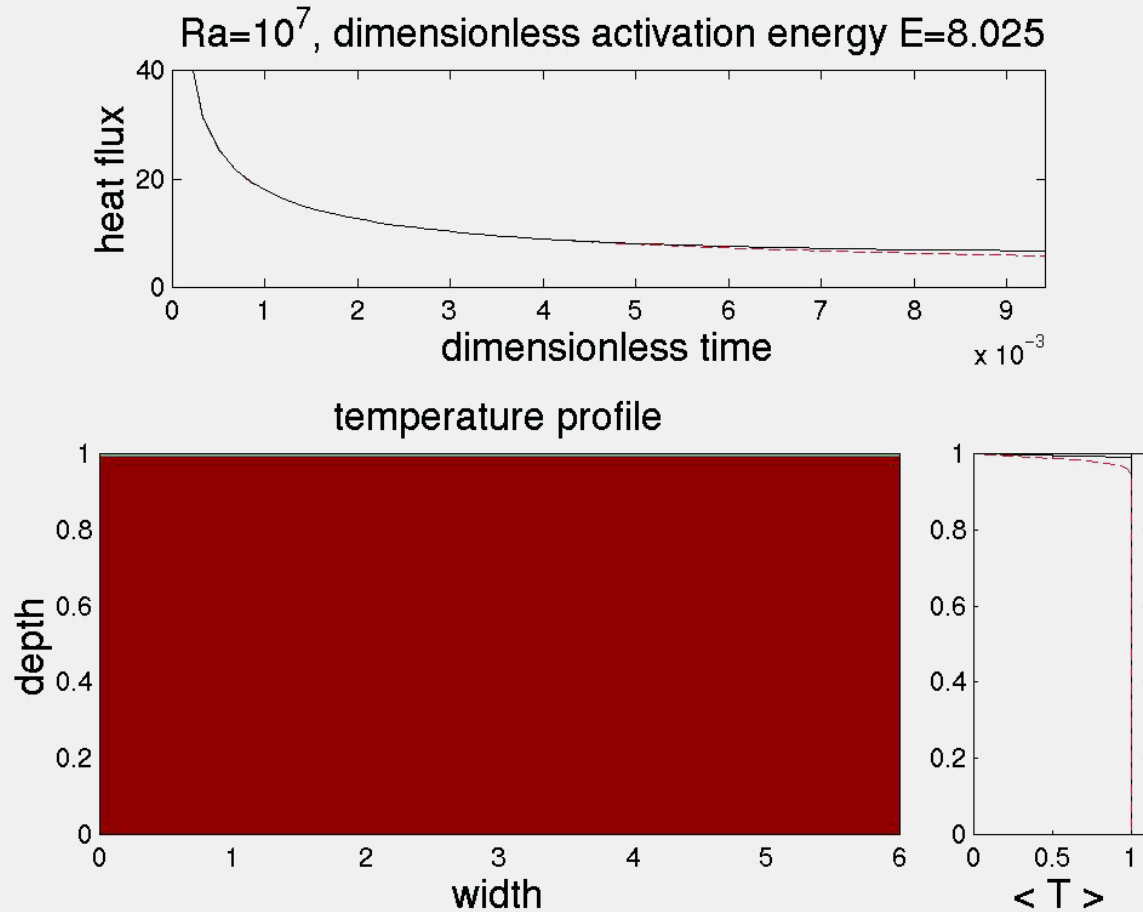
$$\delta \sim (\kappa t)^{1/2}$$

$$\text{Rayleigh number: } Ra = \rho g \alpha (T_m - T_s) \delta^3 / (\kappa \eta)$$

As  $t$  increases,  $\delta$  and  $Ra$  increase. When  $Ra > Ra_{CR}$ , the thermal boundary layer goes unstable and thermal convection starts to happen.



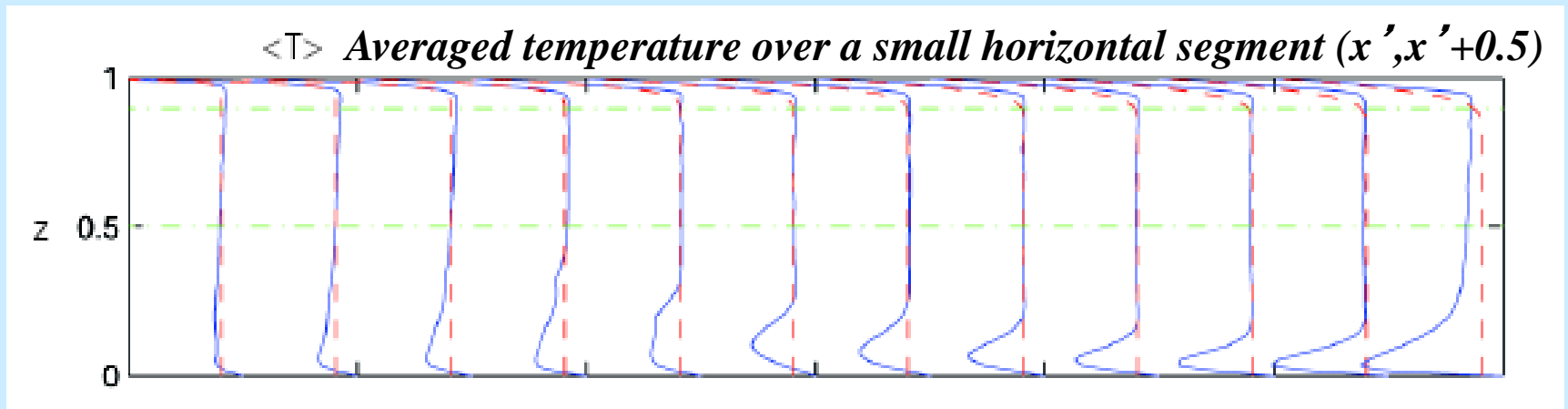
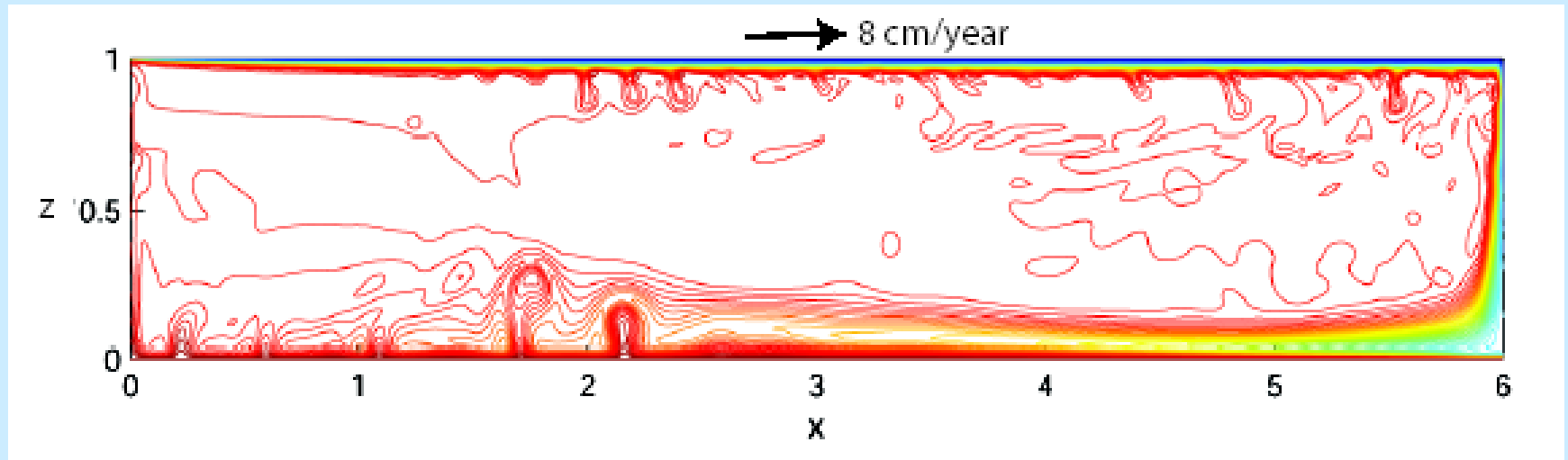
# *SSC from 2D Models (Huang et al., 2003)*



*O’Connell & Hager [1981] pointed out that this vertical exchange of heat from SSC may cause “deepening” not “flattening” in ocean depth.*

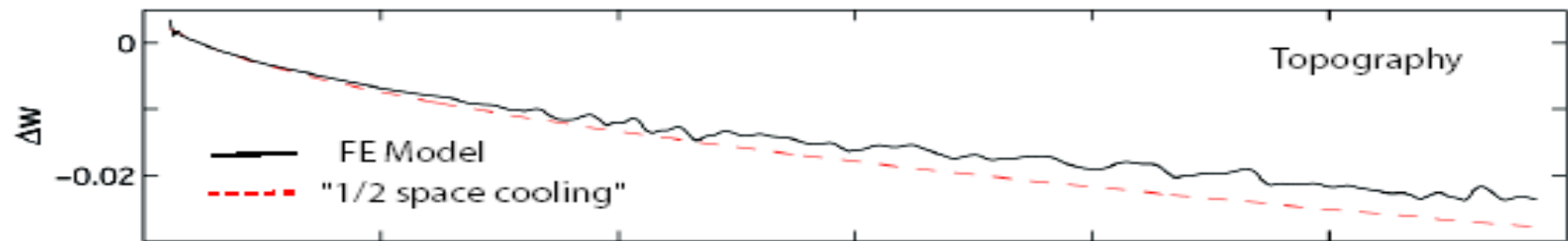
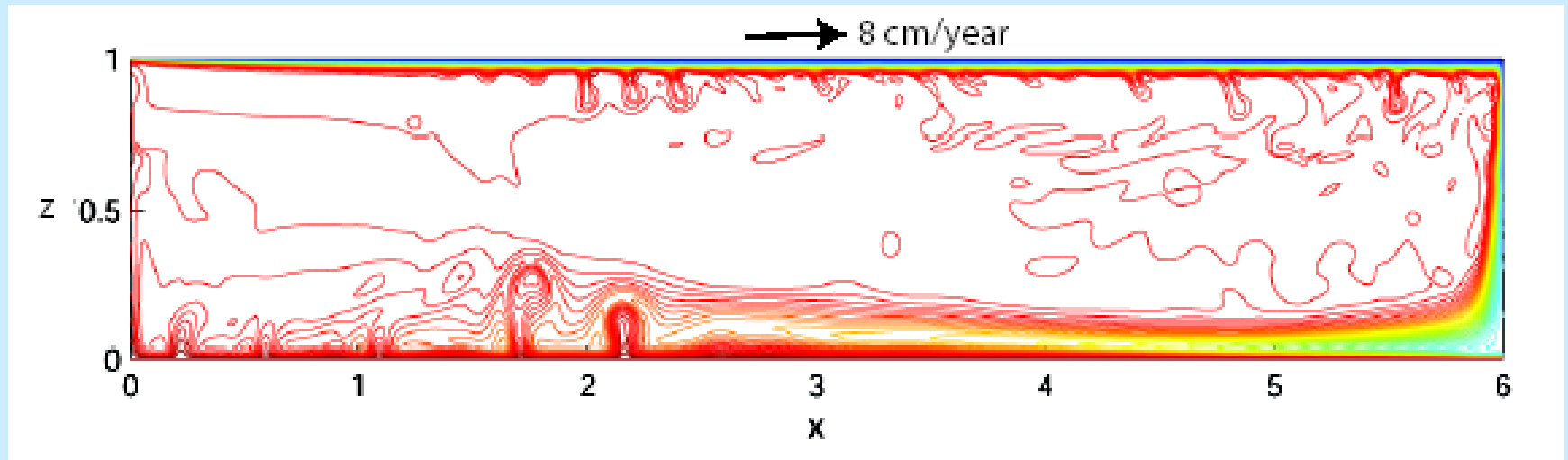
# 2D model with SSC

60% internal heating;  $E=120$  KJ/mol;  $\eta_a=2.6e19$  Pa-s with X60 for the lower mantle

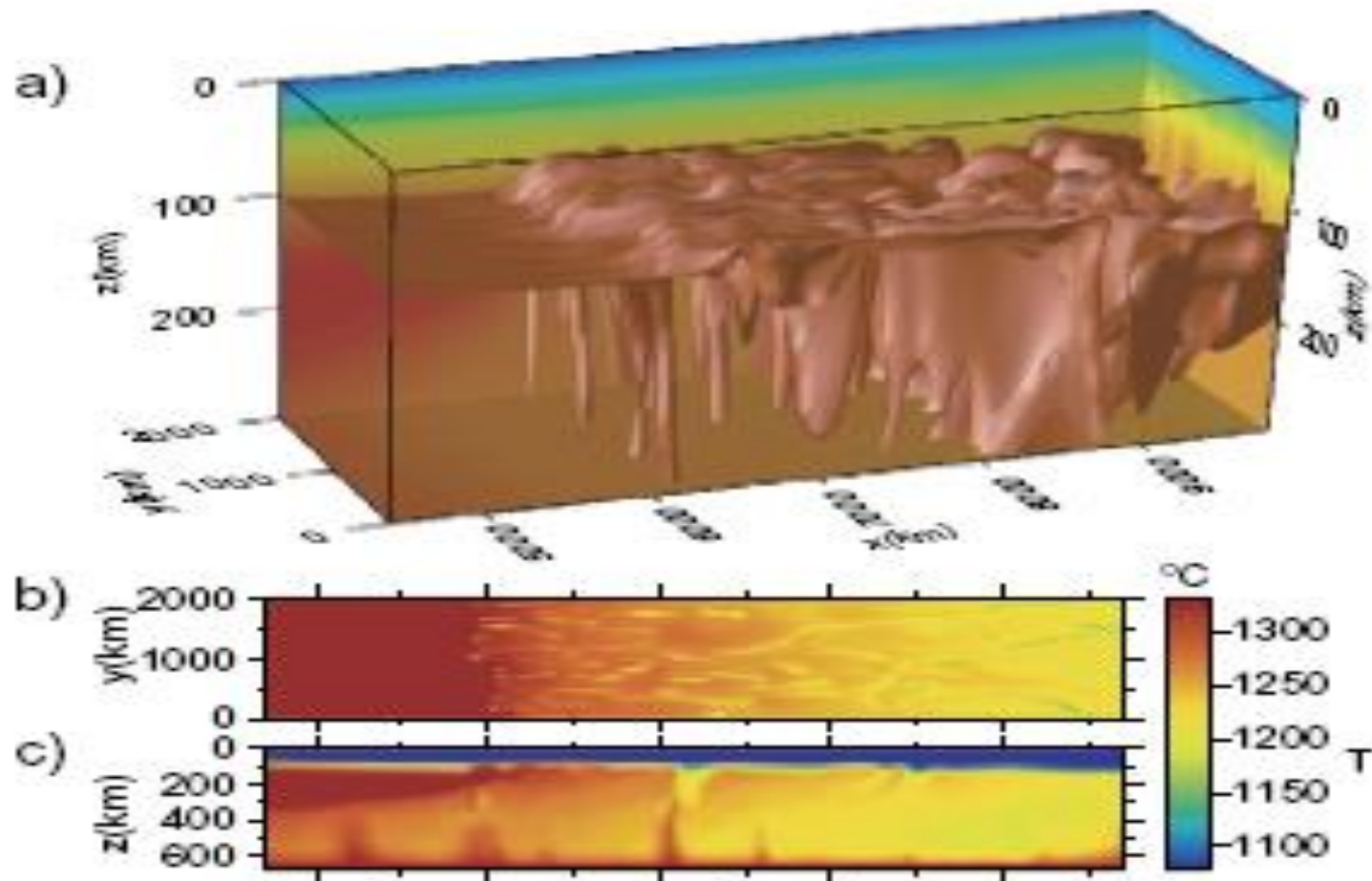


# 2D model with SSC

60% internal heating;  $E=120$  KJ/mol;  $\eta_a=2.6e19$  Pa-s with X60 for the lower mantle



# 3-D modeling of SSC



# ***Summary on lithospheric geotherms***

- ***For oceanic lithosphere: The plate model and  $\frac{1}{2}$  space cooling model are two end-members. The truth is somewhere between.***
- ***For continental lithosphere: Complications with tectonics/volcanism/composition [Jaupart et al., 2007; Lee et al., 2011; Carlson et al., 2005; Rudnick et al., 1999].***

# Convective mantle geotherms

- **Follow adiabatic gradient.**

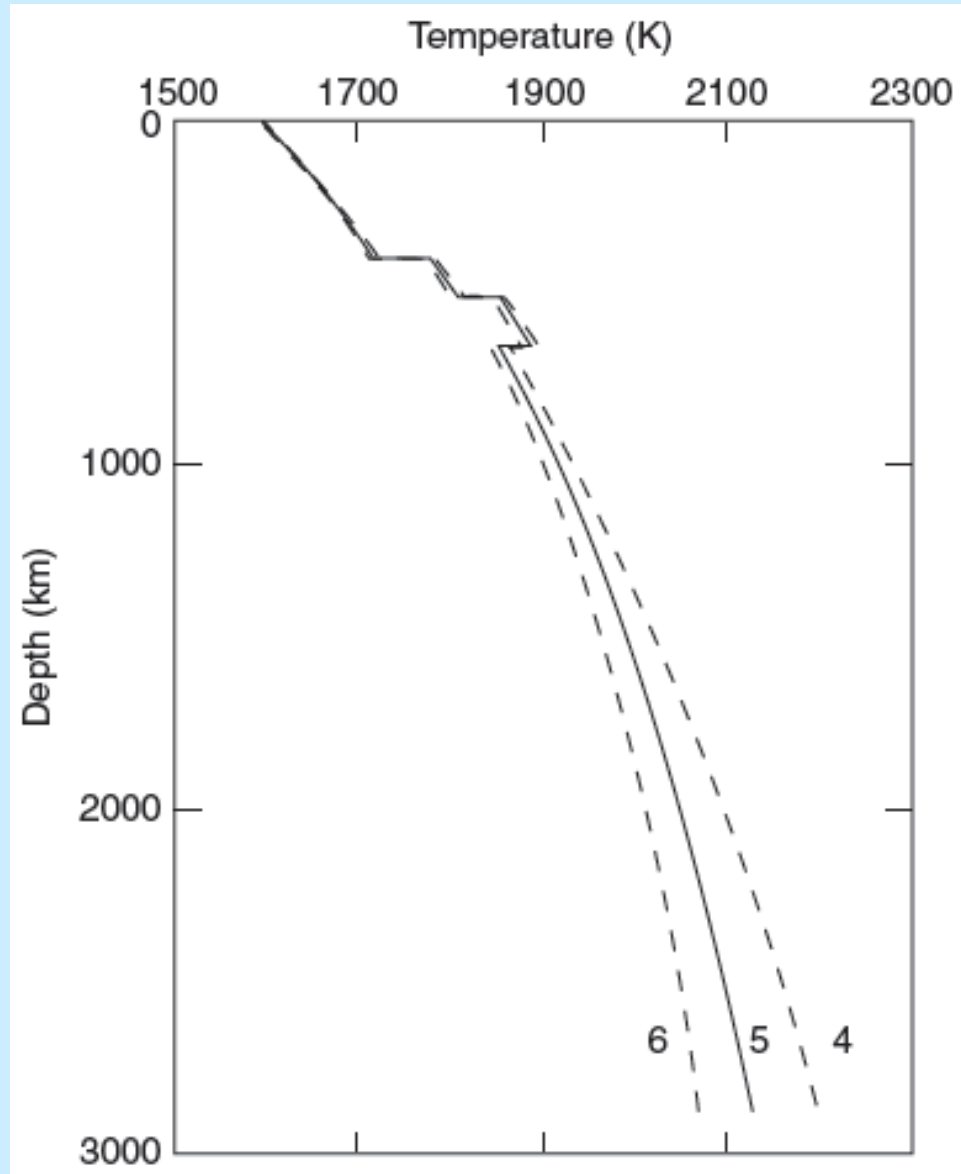
$$(dT/dz)_s = \alpha g T / C_p,$$

**Table 12** Anchor points for the mantle geotherm

Boundary	Depth (km)	Temperature (K)	Reference
MORB generation	50	1590–1750 <sup>a</sup>	Kinzler and Grove (1992)
Olivine–Wadsleyite	410	1760 ± 45	Katsura <i>et al.</i> (2004)
Post-spinel	660	1870 ± 50	Katsura <i>et al.</i> (2003, 2004)
Core–mantle	2884	4080 ± 130	Alfé <i>et al.</i> (2002); Labrosse (2003), this paper

<sup>a</sup>indicates true range of temperatures in the shallow mantle.

***Jaupart et al., 2007***

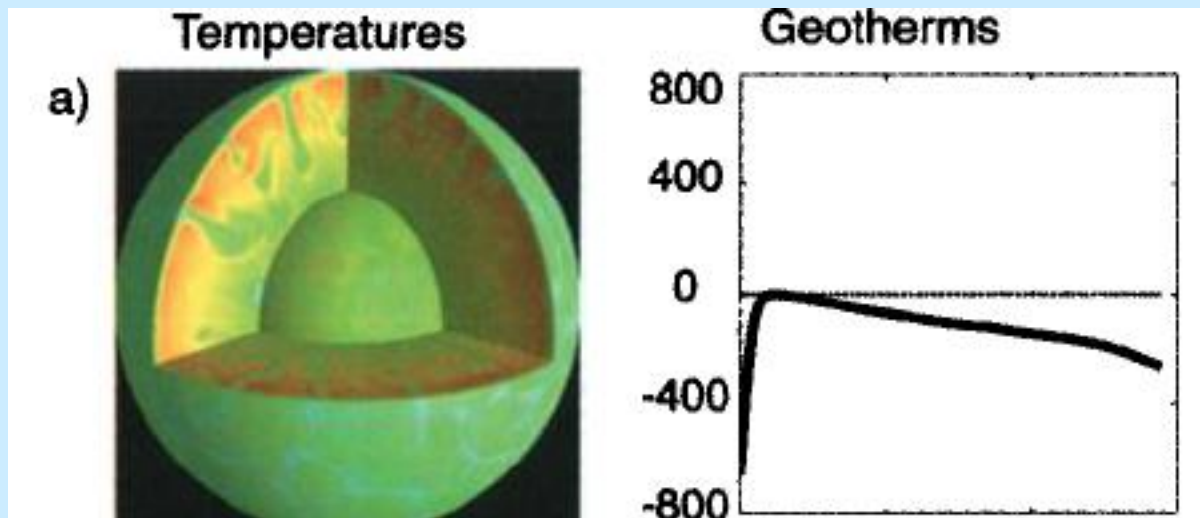


***Katsura et al., 2004; Jaupart et al., 2007***

# ***Sub-adiabatic temperature?***

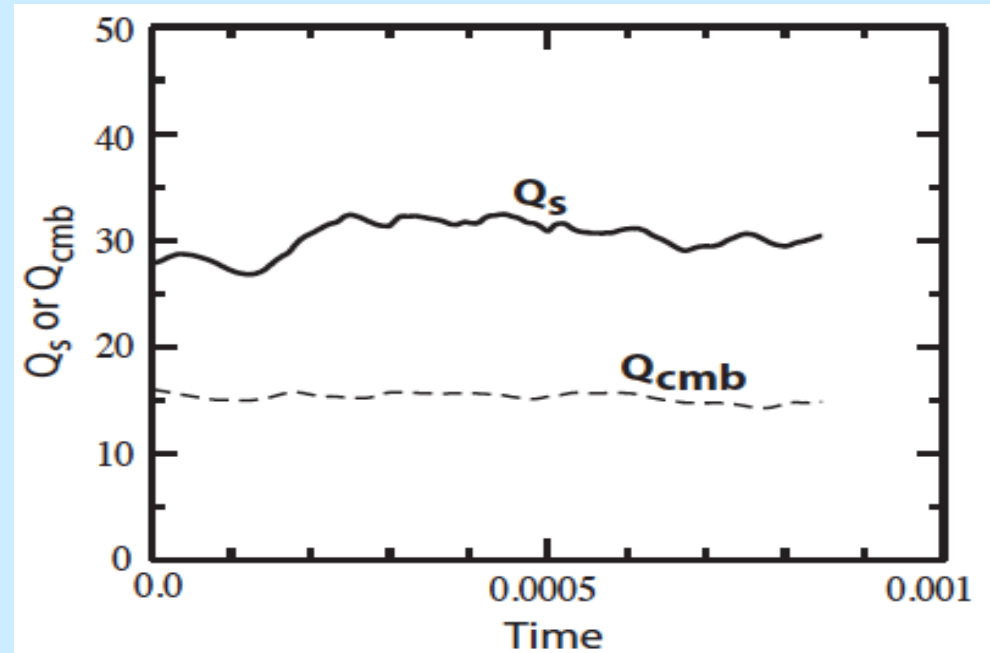
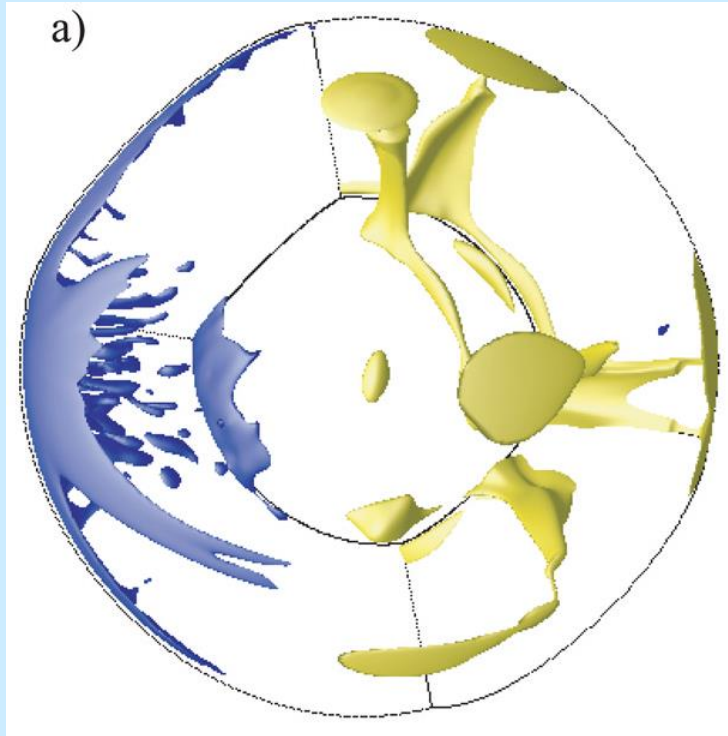
***Jeanoz and Morris [1987] suggested ~400-500 K sub-adiabatic temperature in the lower mantle.***

***Bunge et al. [2001] seemed to support this view.***



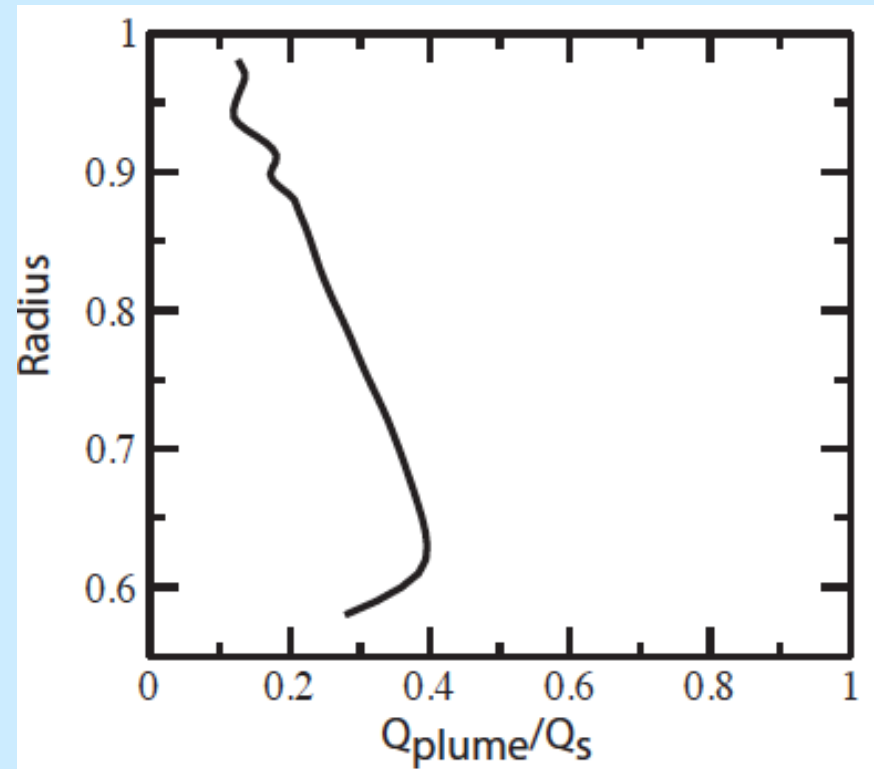
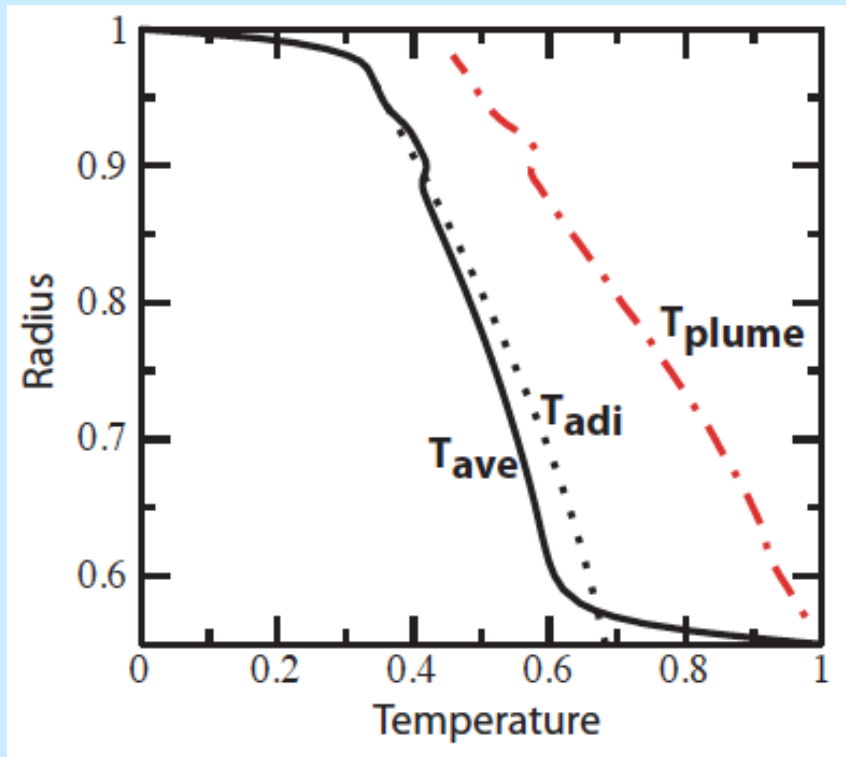


# ***3-D convection model for plume dynamics***



***Zhong, 2006; Leng and Zhong 2008.***

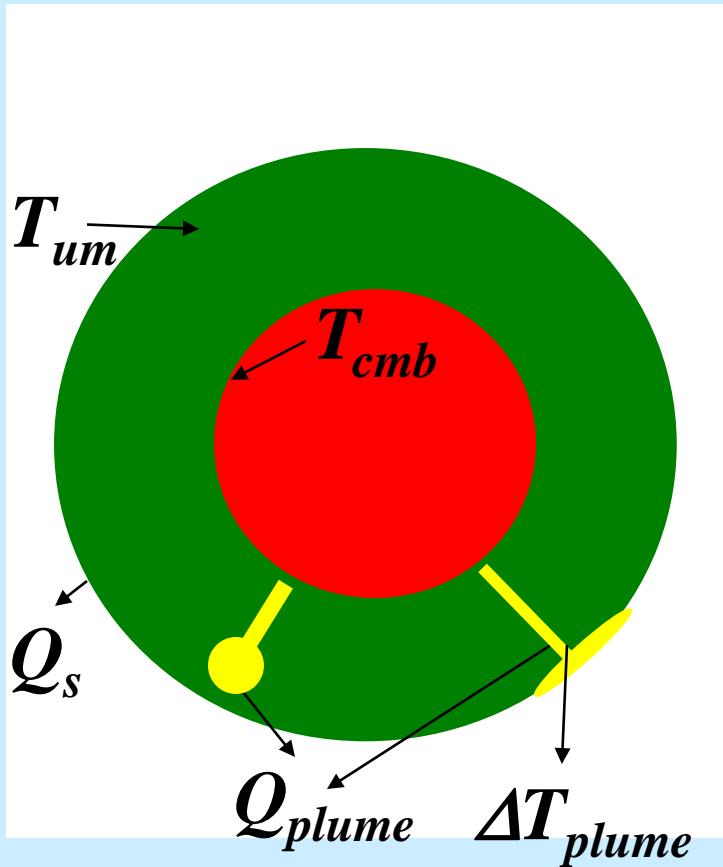
# Temperature and $Q_{\text{plume}}$ vs depths



1) Sub-adiabatic temperature is  $<0.05$  (200 K, for  $T_{\text{cmb}} \sim 4000$  K)

2) Plume excess temperature decreases as plumes rise. This may place constraints on CMB heat flux, using inferred plume flux by Davies [1988] and Sleep [1990].

***Observations in the UM: Plume excess temperature, plume heat flux, and UM and CMB temperature [Zhong, 2006]***

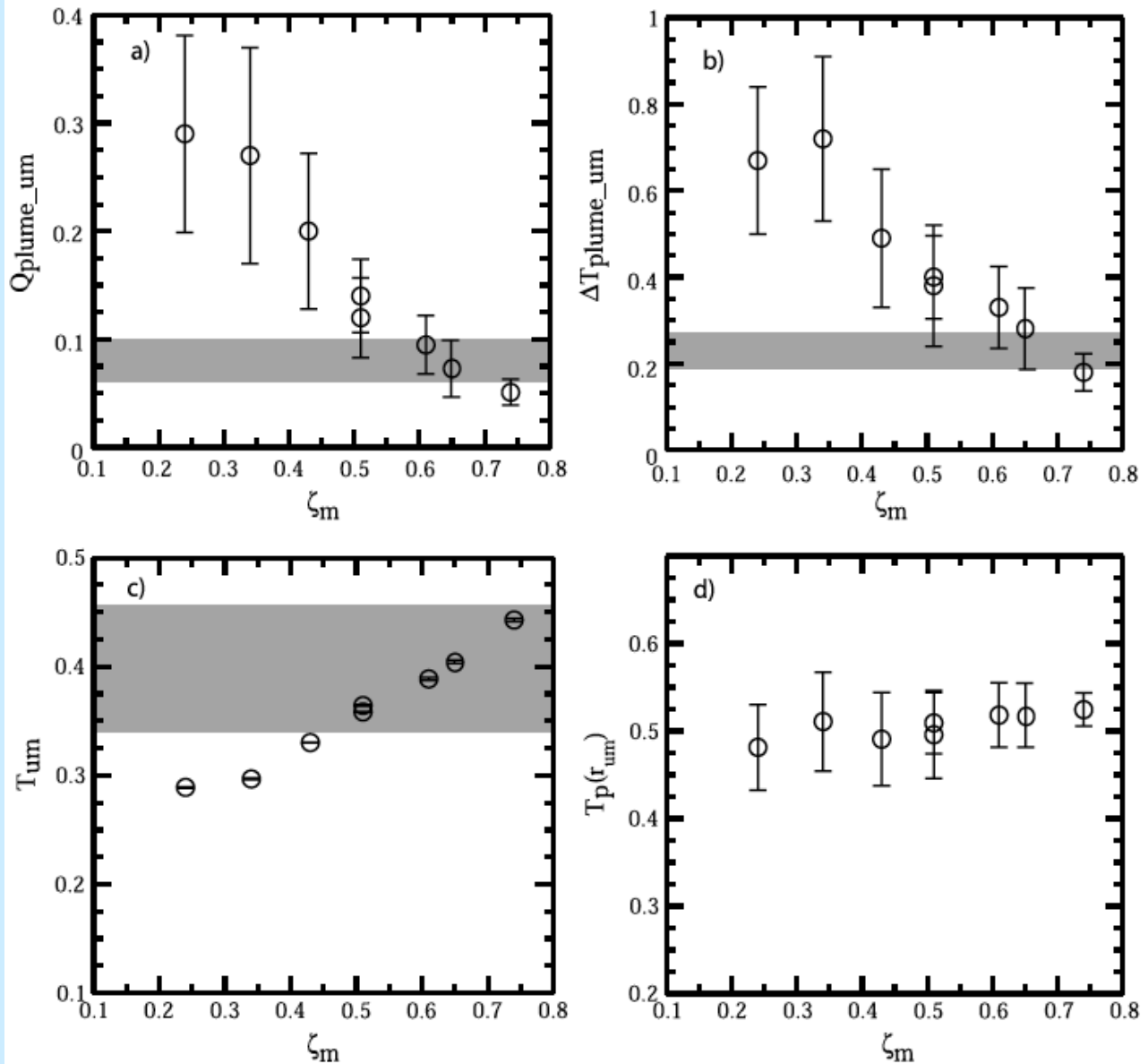


1)  $T_{um} \sim 1280 \text{ }^\circ\text{C}$ ,  $T_{cmb} \sim 3400 \pm 500 \text{ }^\circ\text{C}$   
 $\rightarrow T_{um} / T_{cmb} \sim 0.338-0.455$

2)  $\Delta T_{plume} \sim 250-350 \text{ K}$   
 $\rightarrow \Delta T_{plume} / T_{um} \sim 0.19-0.27$

3)  $Q_{plume} \sim 2.4-3.5 \text{ TW}$   
 $\rightarrow Q_{plume} / Q_s \sim 6-10\%$   
[Davies, 1988; Sleep, 1990].

# Plumes constrain the internal heating ratio to be ~70%

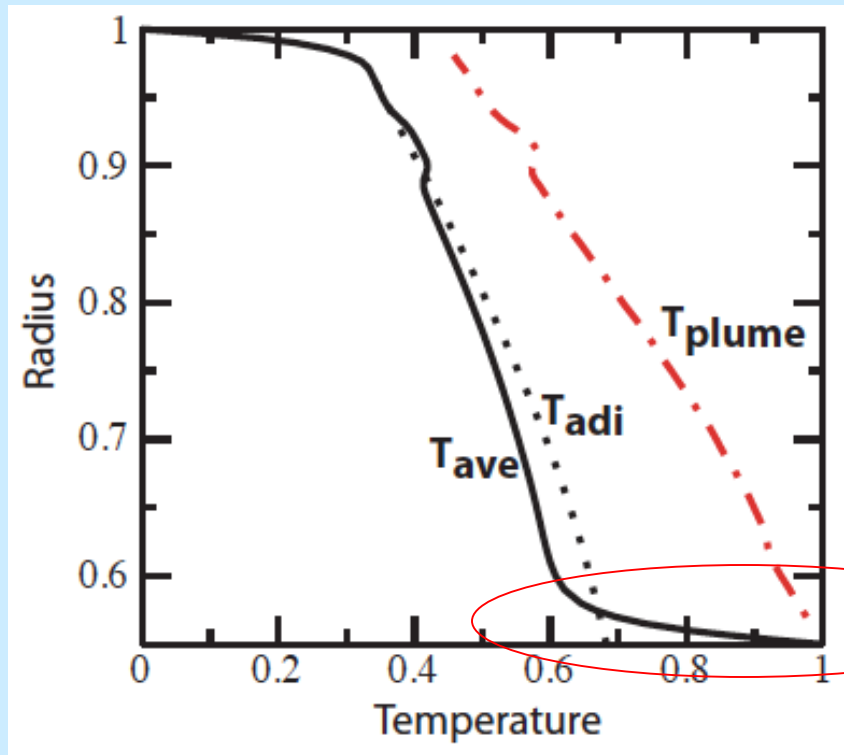


$Q_{\text{cmb}} \sim 30\% Q_s$   
 $\sim 10 \text{ TW}$

# ***Summary on convective mantle geotherms***

- ***Follow adiabats with  $<200$  K subadiabatic temperature in the lower mantle.***
- ***A plume would follow its own adiabat with a steeper slope due to its high temperature (important for some problems, e.g., plume flux).***

# ***Geotherms in the CMB region***

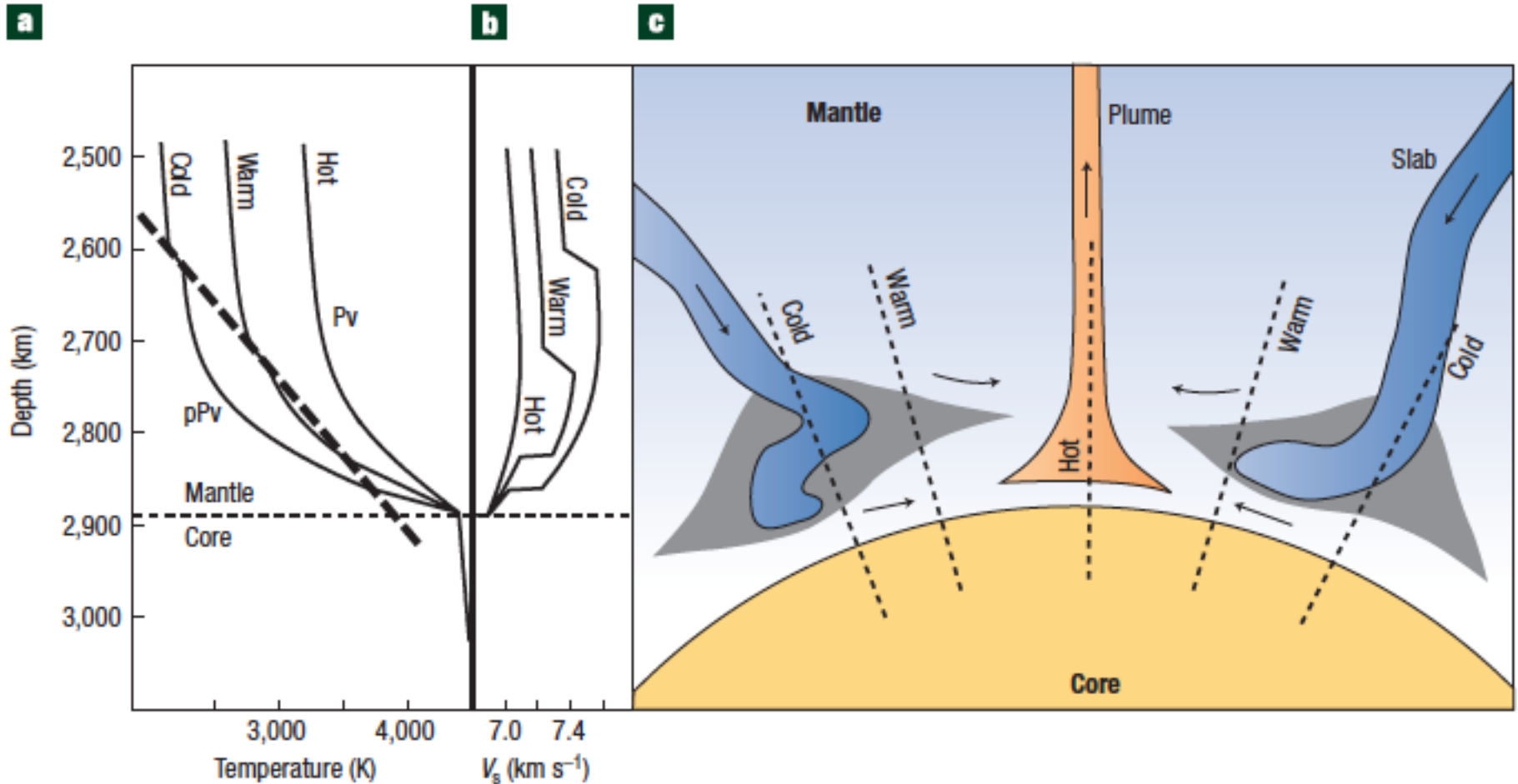


***Another thermal  
boundary layer***

***We need temperature anchors from melting or phase change (e.g., post-perovskite).***

***Also CMB heat flux is useful (from plume estimate, geodynamo, post-perovskite, ...).***

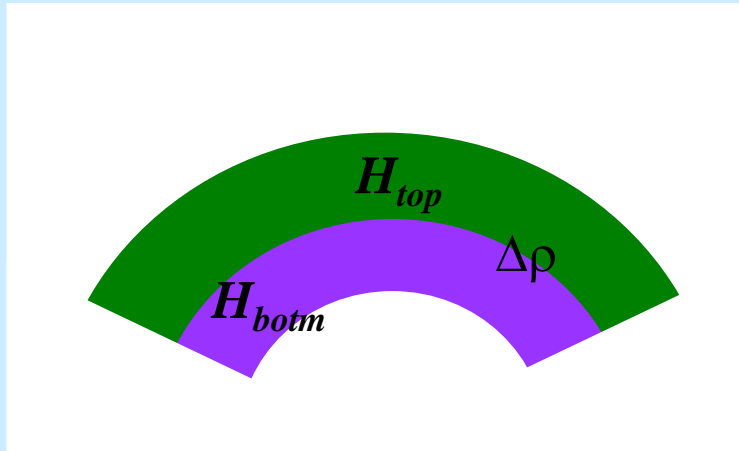
# Post-perovskite phase change



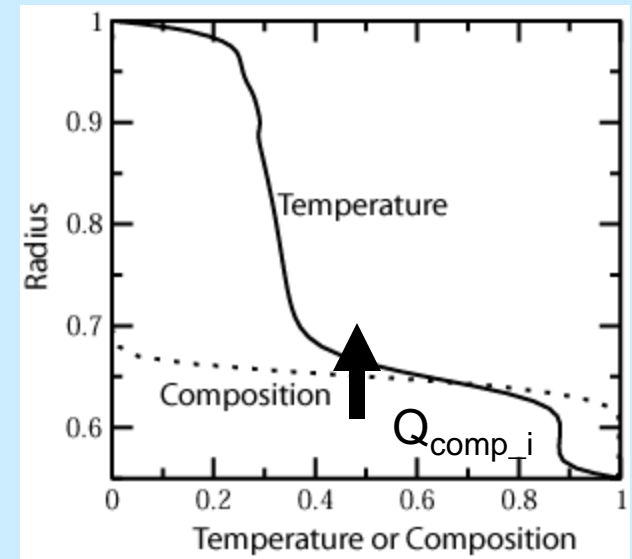
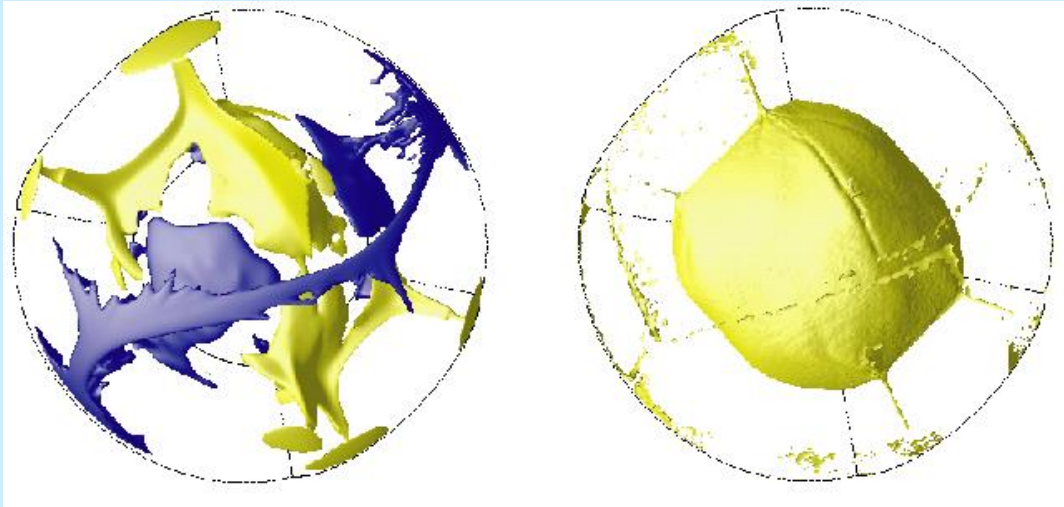
*Lay et al., [2008]; Hernlund et al., [2005]*

*However, post-perovskite phase change may not be a global feature (unlike 410- and 670-km).*

# *Two-layer thermochemical models*



*Thermochemical piles?*



*Yet another thermal boundary layer*





